

VISUAL MODEL BASED SINGLE IMAGE DEHAZING USING ARTIFICIAL BEE COLONY OPTIMIZATION

S.MohamedMansoorRoomi^{#1} R.Bhargavi^{#2} S.Bhumesh^{#3}

Department of Electronics and Communication Engineering, Thiagarajar College of Engineering, Madurai

¹smmroomi@tce.edu

²rbhargavi22@gmail.com

³bhumesh.wave@gmail.com

ABSTRACT

Images are often degraded by atmospheric haze, a phenomenon due to the particles in the air that scatter light. Haze induces a loss of contrast, its visual effect is blurring of distant objects. This paper presents a novel algorithm for improving the visibility of an image degraded by haze. The proposed method uses a cost function based on human visual model to estimate airlight map. It employs Artificial Bee Colony optimization (ABC) as the optimization technique for estimating air light map. Image is dehazed by removing the estimated airlight from the degraded image. The performance of the algorithm is tested and compared with various other dehazing methods and the proposed algorithm dehazes the image effectively outperforming other methods.

KEYWORDS

Airlight map, Artificial Bee Colony optimization, cost function.

1. INTRODUCTION

During propagation solar radiation interacts with the atmosphere, generating a variety effects upon the resulting satellite image that must subsequently be accounted for through atmospheric Corrections. Some atmospheric effects, such as cloud, block almost all radiation in the visible and infrared spectral regions. Others partly obscure the ground-reflected radiation leaving an underlying ground information component in contaminated form. "Haze" is an example of this latter effect. Haze is a commonly used term in image analysis, referring to a set of atmospheric effects that reduce image contrast. In general, the impact of haze is evident when viewing images in blue or green parts of the electromagnetic spectrum. At those wavelengths, it is generally an additive radiometric effect and varies spatially, with the resulting satellite image typically exhibiting the underlying ground cover in a diffused pattern. It is imperative that haze be removed prior to scene analysis. Several different atmospheric scattering or "haze" detection and removal techniques have been reported in the literature. In [7], assuming the scene depth is given, atmospheric effects are removed from terrain images taken by a forward-looking airborne camera. In [4] when two images are given, polarised haze effects are removed. For each photographic image, a polarising filter is attached where the image acquisition device is identically placed. The resulting images would be images which differ only in magnitude of polarised haze light

component. In [9] estimation of several parameters are done automatically with the assumption that the higher spatial-bands of the direct transmission, the surface radiance reaching the camera, and the polarized haze contribution are uncorrelated. In [12] depending on transmission and to suppress noise amplification during dehazing, a regularisation mechanism is proposed. A user interactive tool for removing weather effects is described in [2] This method requires manually to identify regions that are heavily affected by haze or to provide some coarse depth information. In the airlight is assumed to be constant over the entire image and is estimated given a single image, based on the fact that in natural images the local sample mean of pixel intensities is proportional to the standard deviation. This work presents a novel image processing technique to estimate optimal airlight map as in Section II. Results and Conclusion are provided in Section III and Section IV.

2. PROPOSED METHOD

The first phase of this work involves estimation of optimal air light map using Artificial Bee Colony Optimization(ABC) second phase involves an effective method to correct the degraded image by subtracting the estimated air light map from the degraded image. The flowchart for proposed algorithm is shown in Figure 1.

2.1. Haze Effect on Image

The hazy image is degraded by airlight that is caused by scattering of light with particles in air. Airlight plays the role of being an additional source of light as modeled in [6] and Eqn (1) below.

$$I'_{R,G,B} = I_{R,G,B} + \lambda_{R,G,B} \quad (1)$$

where $I'_{R,G,B}$ is the degraded image, $I_{R,G,B}$ is the original image, and $\lambda_{R,G,B}$ represents the airlight for the Red, Green and Blue channels. This relationship can be applied in the case where airlight is uniform throughout the whole image. However, the contribution of airlight is not usually uniform over the image because it is a function of the visual depth, which is the distance between the camera and the object. Therefore, the model can be modified to reflect the depth dependence as follows.

$$I'_{R,G,B}(d) = I_{R,G,B}(d) + \lambda_{R,G,B}(d) \quad (2)$$

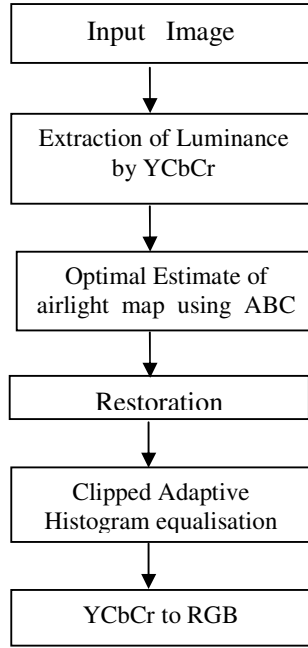


Figure 1 .Flowchart of the Proposed Algorithm

Note that “d” represents depth. Unfortunately, it is very difficult to estimate the depth using one image taken in hazy weather conditions, so we present an airlight map that models the relationship between the coordinates of the image pixels and the airlight. In this paper, since the amount of scattering of a visible ray by large particles like haze and clouds are almost identical, the luminance component is used alone to estimate the airlight instead of estimating the R, G, and B components. The luminance image can be obtained by a fusion of the R, G, and B components. Subsequently, the color space is transformed from RGB to YCbCr. Therefore Eqn (2) can be re-expressed as follows.

$$Y'_{R,G,B}(d) = Y_{R,G,B}(d) + \lambda_{R,G,B}(d) \quad (3)$$

where Y' and Y reflect the degraded luminance and clear luminance images respectively at position (i,j). In order to restore the image blurred by haze, we need to estimate the airlight map and subtract the airlight from the hazy image as follows.

$$\hat{Y}(i, j) = Y'(i, j) - \hat{\lambda}_s(i, j) \quad (4)$$

In this model, \hat{Y} represents the restored image and $\hat{\lambda}$ is the estimated airlight map.

2.2. Optimal Estimate of Airlight Using Artificial Bee Colony

The luminance image is divided into various blocks and for each block airlight is estimated. In order to estimate the air light, we improved the cost function method in [6] using a compensation that is based on the human visual model. In Eqn (3), the airlight is to be estimated to restore the image degraded by haze. To estimate the airlight, the human visual model is employed. As described by Weber's law, a human is more insensitive to variations of luminance in bright regions than in dark regions.

$$\Delta S = k \frac{\Delta R}{R} \tag{5}$$

where R is an initial stimulus, ΔR is the variation of the stimulus, and ΔS is a variation of sensation. In the hazy weather conditions, when the luminance is already high, a human is insensitive to variations in the luminance. We can estimate the existing stimulus in the image signal by the mean of the luminance within a region. The variation between this and hazy stimulus can be estimated by the standard deviation within the region. Thus the human visual model would estimate the variation of sensation as

$$\frac{STD(Y)}{mean(Y)} = \frac{\sqrt{\frac{1}{n} \sum (y_i - \bar{Y})^2}}{\bar{Y}} \tag{6}$$

$$A(\lambda) = \frac{STD(Y' - \lambda)}{mean(Y' - \lambda)} \tag{7}$$

In Eqn (7), increasing λ causes an increase in $A(\lambda)$, which means that a human can perceive the variation in the luminance. However, if the absolute value of the luminance is too small, it is not only too dark, but the human visual sense also becomes insensitive to the variations in the luminance that still exist. To compensate for this, a second function is generated as follows.

$$B'(\lambda) = mean(Y' - \lambda) \tag{8}$$

Eqn (8) indicates information about mean of luminance. In a hazy image, the result of Eqn (8) is relatively large. And, increasing λ causes a decrease in $B'(\lambda)$ which means that overall brightness of the image decreases. Functions (7) and (8) reflect different scales from each other. Function (8) is re-scaled to produce Eqn (9) to set 0 when input image is Ideal. Note that "Ideal" represents the ideal image having a uniform distribution from the minimum to the maximum of the luminance range. In general, the maximum value is 235 while the minimum value is 16.

$$B(\lambda) = mean(Y') - \lambda * \frac{STD(ideal)}{mean(ideal)} \tag{9}$$

The λ satisfying Eqn (10) is the estimated optimal airlight.

$$\hat{\lambda} = \arg \min_{\lambda} \{ | A(\lambda) - B(\lambda) | \} \tag{10}$$

Estimation of optimal air light is performed by Artificial Bee Colony Optimization. In the ABC algorithm, the position of a food source represents a possible solution of the optimization problem and the nectar amount of a food source corresponds to the quality (fitness) of the associated

solution. The number of the employed bees or the onlooker bees is equal to the number of solutions in the population. At the first step, the ABC generates a randomly distributed initial population $P(G = 0)$ of S_N solutions (food source positions), where S_N denotes the size of population. Each solution (food source) $x_i (i = 1, 2, \dots, S_N)$ is a D -dimensional vector. Here, D is the number of optimization parameters. The population of the positions (solutions) is subjected to repeated cycles, $C = 1, 2, \dots, C_{max}$, of the search processes of the employed bees, the onlooker bees and scout bees when the initialization is completed. In the memory of onlooker bee, a position for finding a new food source and to test the nectar amount (fitness value) of the new source (new solution) is produced. If the nectar amount of the new source is higher than that of the previous one the bee remembers the new position and eliminates the previous position else the previous position is maintained. When the search process is completed by all employed bees, the nectar information of the food sources (solutions) is shared among them and their position information with the onlooker bees on the dance area. An onlooker bee evaluates the nectar information taken from all employed bees and chooses a food source with a probability related to its nectar amount. As in the case of the employed bee, she produces a modification on the position (solution) in her memory and checks the nectar amount of the candidate source (solution). Providing that its nectar is higher than that of the previous one, the bee memorizes the new position and forgets the old one.

An onlooker bee chooses a food source depending on the probability value associated with that food source, p_i , calculated by the following expression (11)

$$P_i = \frac{FIT_i}{\sum_{n=1}^{S_N} FIT_n} \quad (11)$$

where fit_i is the fitness value of the solution i evaluated by its employed bee, which is proportional to the nectar amount of the food source in the position i and S_N is the number of food sources which is equal to the number of employed bees (B_N). In this way, the employed bees exchange their information with the onlookers. In order to produce a candidate food position from the old one, the ABC uses the following expression (12)

$$v_{ij} = x_{ij} + \varphi_{ij}(x_{ij} - x_{kj}) \quad (12)$$

Where $k \in \{1, 2, \dots, B_N\}$ and $j \in \{1, 2, \dots, D\}$ are randomly chosen indexes. Although k is determined randomly, it has to be different from i . $\varphi_{i,j}$ is a random number between $[-1, 1]$. It controls the production of a neighbour food source position around $x_{i,j}$ and the modification represents the comparison of the neighbour food positions visually by the bee. Equation 12 shows that as the difference between the parameters of the $x_{i,j}$ and $x_{k,j}$ decreases, the perturbation on the position $x_{i,j}$ decreases, too. Thus, as the search approaches to the optimum solution in the search space, the step length is adaptively reduced.

2.3. Smoothed Airlight Map

When the airlight map obtained through ABC is directly used for restoration, it would create blocking lines effects in the restored image. The sharp transitions among the blocks creates undesired effects. In order to resolve this issue, the airlight map is further processed. The resultant

smoothed airlight map $\hat{\lambda}_s$ is generated by further convolving optimal airlight map

($\hat{\lambda}$) with Gaussian window which is used to restore haze free image.

2.4. Restoration

In order to restore the luminance image, the smoothed optimal airlightmap($\hat{\lambda}_s$) is subtracted from the degraded image as Eqn (13). The term 'b' is mean value of airlight map obtained.

$$\hat{Y}(i, j) = Y'(i, j) - \hat{\lambda}_s(i, j) + b \quad (13)$$

The restored luminance image (\hat{Y}) is further processed by clipped adaptive histogram equalization.

2.5. Clipped Adaptive Histogram Equalization

CLAHE addresses the issue of over amplification of noise that adaptive histogram equalisation produces by limiting the contrast enhancement. The slope of the transformation function gives the contrast amplification in the vicinity of a given pixel value. This is proportional to the slope of the neighbourhood cumulative distribution function (CDF) and therefore to the value of the histogram at that pixel value. Before computing the CDF, amplification is limited by clipping the histogram at a predefined value thereby limiting the slope of the CDF and therefore of the transformation function. The value at which the histogram is clipped, the so-called clip limit, depends on the normalization of the histogram and thereby on the size of the neighbourhood region. CLAHE operates on small regions in the image, called tiles, rather than the entire image. Each tile's contrast is enhanced, so that the histogram of the output region approximately matches the histogram specified by the 'Distribution' parameter. The neighbouring tiles are then combined using bilinear interpolation to eliminate artificially induced boundaries. The contrast, especially in homogeneous areas, can be limited to avoid amplifying any noise that might be present in the image. By changing the colour space from YCbCr to RGB, dehazed colour image can be obtained.

3. RESULTS AND DISCUSSION

The proposed algorithm is tested with haze images and compared against various Dehazing algorithms. He et al proposed Dehazing based on dark channel prior [14], Ding et al proposed Colour contrast enhancement [15], Xie et al proposed haze removal based on retinex theory [16]. The proposed method is also compared with dehazing using Particle Swarm optimisation technique [17]. The effectiveness of the proposed algorithm is evaluated by the tenengrad criterion and contrast improvement index.

3.1 Tenengrad

In order to evaluate the effectiveness of the resultant image a well-known benchmark-image sharpness measure, the tenengrad criterion as in Eqns (14)-(15) are used. The tenengrad criterion is based on gradient, at each pixel(x, y), where the partial derivatives are obtained by a high-pass filter, eg., The gradient magnitude is given by: sobel operator, with the convolution kernels

$$S(x, y) = \sqrt{(i_x * I(x, y))^2 + (i_y * I(x, y))^2} \quad (14)$$

$$TEN = \sum_x \sum_y S(x, y)^2, S(x, y) > T \quad (15)$$

And the tenengrad criteria is formulated as were T is the threshold. The quality of the image is usually considered better if its tenengrad value is higher.

3.2 Contrast Improvement Index

Let I be the given image of size NxM. For every pixel at I(x,y) a block of n1 x n1 size is taken. Each block is passed to the contrast finding function. Each block passed is further broken into small windows of size n2 x n2. Mean pixel intensity M(w) is calculated for each of the windows using the equation (16).

$$M(w) = \frac{\sum_{x=0}^W \sum_{y=0}^W b(x,y)}{W_x W_y} \quad (16)$$

N-no of windows in the block

W-size of the window

Further the Mean absolute deviation D(w) is calculated for each window.

$$D(w) = \frac{\sum_{x=0}^W \sum_{y=0}^W b(x,y) - M(w)}{W_x W_y} \quad (17)$$

Using the calculated Mean pixel intensity, M(w) and Mean absolute deviation, D(w) values the Deviation of the mean pixel intensity, D_m and the Mean of the deviation values μ_d are calculated.

$$\mu_d = \frac{\sum_{y=0}^W D(w)}{N} \quad (18)$$

$$D_m = \frac{\sum_{x=0}^W M(w) - \sum_{y=0}^W \frac{D(w)}{N}}{N} \quad (19)$$

The sum of the calculated deviation and calculated mean would represent the contrast of the nth block, C_n . Similarly the contrast of each and every block would be calculated. Contrast value of the nth block is given by Eqn (20)

$$C_n = \mu_d + D_m \quad (20)$$

B-total no of blocks. In this way the contrast of the image is calculated. Improved contrast measure shows the significance of the proposed method.

The tenengrad criterion is based on gradient at each pixel and the image quality is usually considered higher if its tenengrad value is larger. Results are shown in Figure 2-4 and the corresponding tenengrad values and contrast improvement index for various images is computed and tabulated in Table.1 and Table.2. Figure 2(a) and Figure 2(b) shows the restored image by estimated airlight map of image 1 before and after smoothing respectively. Figure 3 and Figure 4 shows the dehazed images obtained by various other Dehazing algorithms.



Figure 2 (a) Estimated Airlight map Restored Image before Smoothing



Figure 2 (b) Estimated Airlight map Restored Image after Smoothing

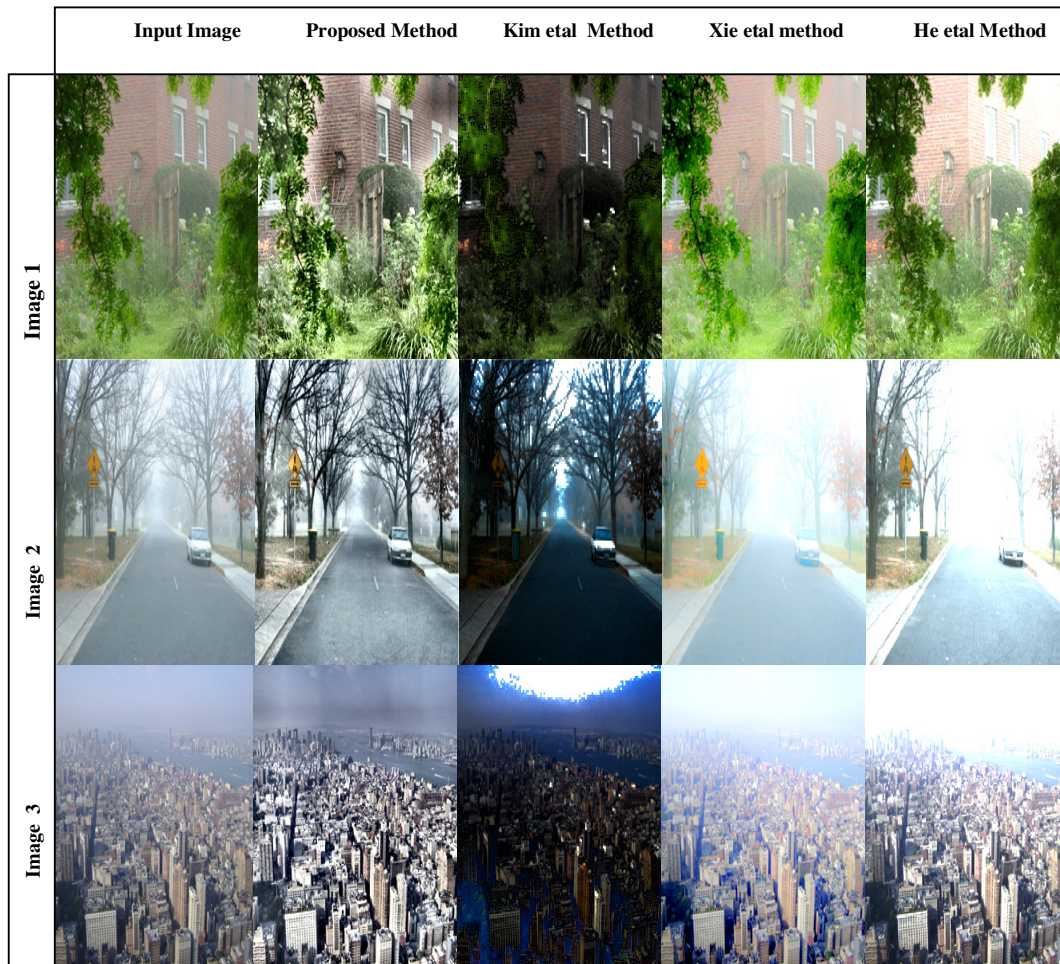


Figure 3 Hazy Images as Input and Comparison of Dehazed Images by various algorithms

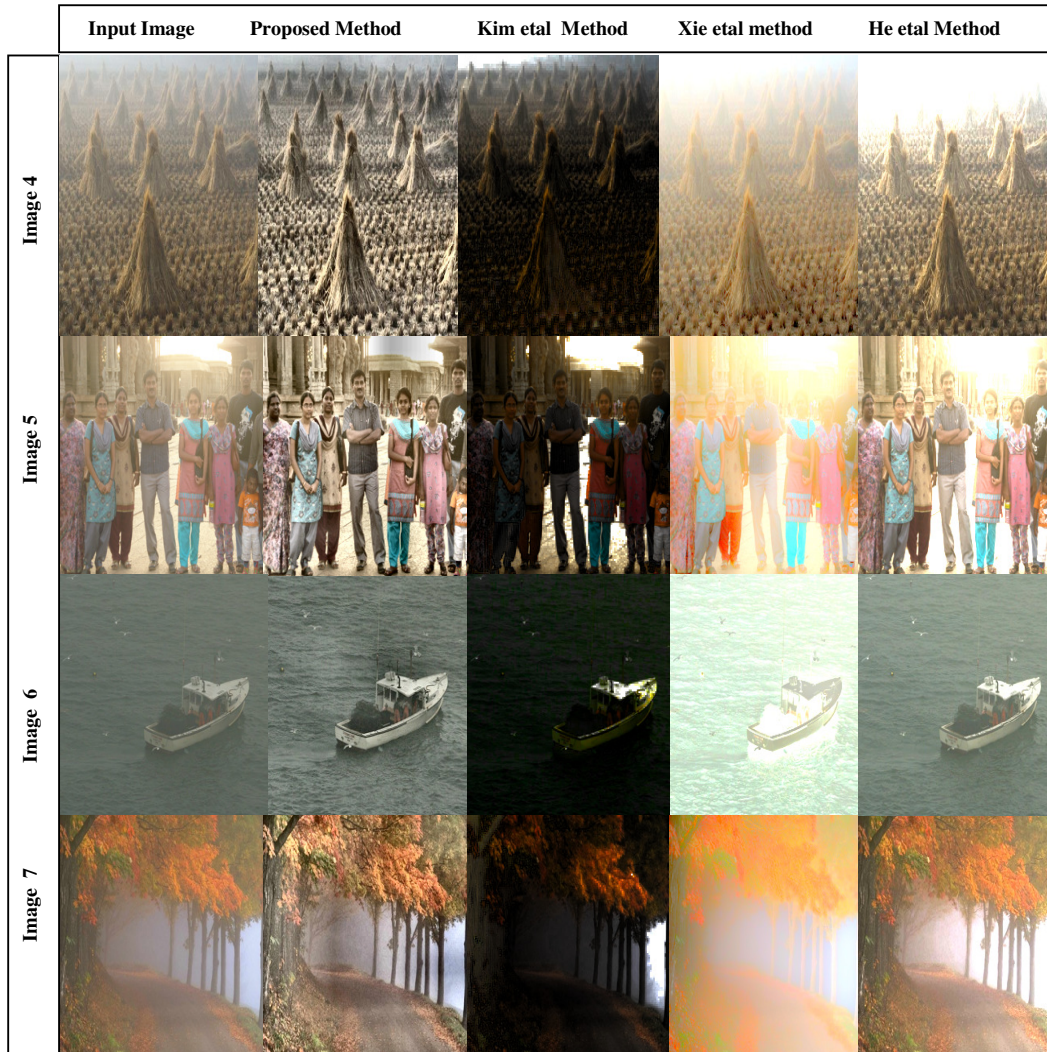


Figure 4 Hazy Images as Input and Comparison of Dehazed Images by various algorithms

Table 1 and Table 2 depicts the proposed algorithm has better performance having the highest tenengrad and contrast improvement index measures among other methods. Figure 3 and Figure 4 depicts various results of the algorithms wherein the other methods does not remove haze completely or degrade the image resulting in poor quality. The proposed method preserves the details of the image and also removes the haze efficiently.

Table 1.Tenengrad values

Image	Original image	He etal method	Ding etal method	Xie etal method	Dehazing using PSO	Proposed method using ABC
1	1309831	92049	3031144	788740	7319721	52092818
2	701286	30631	974296	12460000	8703646	46385527
3	1582132	100053	3878440	1274500	14602661	54763393
4	613077	107355	894088	547420	5707143	54094743
5	659622	74909	1544002	5349100	8523587	53026664
6	605532	64051	965403	431900	707074	15311663
7	548885	66246	1479296	1249900	2819934	34701134

Table 2.Contrast Improvement Index

Image	Original image	He etal method	Ding etal method	Xie etal method	Dehazing using PSO	Proposed method using ABC
1	39.3199	0.2630	53.2344	0.1164	40.9097	38.6072
2	44.8933	0.2489	38.8303	45.1062	67.3529	68.4652
3	28.904	0.2121	39.1314	11.4662	44.0511	44.2775
4	36.3144	0.2358	33.8525	40.6342	44.3183	46.6907
5	28.3586	0.1679	13.9295	48.5689	39.5221	45.3535
6	8.9947	0.0485	17.3126	0.0469	24.2855	26.8489
7	38.2544	0.1926	42.2278	0.0687	50.9545	52.2699

4. CONCLUSION

Optimal airlight is estimated using ABC and dehazed image is restored by a novel technique based on the human visual model which has been proposed in this paper. Artificial Bee Colony Optimisation performs better than Particle Swarm Optimisation in dehazing and the contrast measures are found to be increased. The performance of the proposed work is validated quantitatively with contrast parameter Tenengrad and contrast improvement index.

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Authors

S.Mohamed Mansoor Roomi received his B.E degree in Electronics and communication Engineering from Madurai Kamarajar University, in 1990.and the M.E(Power Systems)&ME (Communication Systems) from Thiagarajar College of Engineering, Madurai in 1992&1997 and PhD in 2009 from Madurai Kamarajar University . His primary Research Interests include Image Enhancement and Analysis.



R.Bhargavi is currently doing her under graduation in Electronics and Communication Engineering at Thiagarajar College of Engineering, Madurai,India. Her research interests are in the area of Image Enhancement and Image Analysis.



S.Bhumesh is currently doing his under graduation in Electronics and Communication Engineering at Thiagarajar College of Engineering, Madurai,India. His research interests are in the area of Image Enhancement and Image Analysis.

